Searching for Unassociated Fermi Objects with the highest energy gamma rays in the Galaxy

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Outline

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- Cosmic and Gamma-Ray Astronomy
- HAWC and Fermi-LAT Observatories
- The 3HWC and 4FGL Catalogs
- Method
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- Conclusion and Outlook

Cosmic Rays

- Cosmic rays are atomic nuclei constantly bombarding us from space
- Sources include supernova explosions and super-massive black holes
- Observations of cosmic rays above 10^20 eV





black holes

Gamma rays

They point to their sources, but they can be absorbed and are created by multiple emission mechanisms.

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Earth

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air shower

Neutrinos

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They are weak, neutral particles that point to their sources and carry information from deep within their origins.

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Cosmic rays They are charged particles and are deflected by magnetic fields.

Energy ranges for photons



Energy ranges for photons



Fermi-LAT Satellite

- The Large Area Telescope (LAT) detects gamma rays in the lower-energy range of 20 MeV to 300 GeV (~10^6 eV to ~10^11 eV)
- Instantaneous field of view is ~2 steradians
- Has 36 layers of silicone strip detectors and 16 layers of tungsten foil



Fermi-LAT's Field-of-view



Detection techniques

- Fermi-LAT has no sensitivity to rays with higher energies due to limited detection size
- Use indirect detection with ground based detectors to observe high-energy gamma rays



Detecting Extensive Air Showers



The observation of an extensive air shower particle cascade using by collecting Cherenkov radiation (left) and observing the shower particles at ground level (right). From the Milagro collaboration (copyright © 2002, University of California).

The High Altitude Water Cherenkov (HAWC) Observatory

- 300 water tanks, each with ~200,000 liters of water
- Observes up to two-thirds of our sky
- Detects gamma-ray showers through the Water Cherenkov method
- Energy range of ~10^11 eV to ~10^14 eV
- Gamma ray signal is along direction of initial particle; cosmic ray signal breaks apart



Simulation of a gamma-ray induced shower



Run 2118, TS 45004, Ev# 41, CXPE40= 55.7, Cmptness= 10.7



HAWC's Field-of-view





The 3HWC and 4FGL Catalogs

- 3HWC Catalog
- 1,523 days of HAWC data
- Data from 65 TeV candidate sources
 - 20 unassociated sources

- 4FGL Catalog
- 10 years of Fermi-LAT data
- Data from 5,064 GeV sources
- 1,410 unassociated sources

Associated and Unassociated Sources in the 3HWC catalog



Associated and Unassociated Sources in the 3HWC catalog



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Associated and Unassociated Sources in the 3HWC catalog



The goal: to find spatial associations between the unassociated sources from HAWC and Fermi

Method

Criteria for Investigating 3HWC Sources

- Must have unassociated Fermi sources within a distance of 1° nearby
- Fermi sources are plotted with their 3HWC counterparts on significance maps

3HWC J1847-017



Results – 14 3HWC Sources

3HWC J1843-034 \rightarrow J1843.7-0328



3HWC J2023+124 ext. \rightarrow J2024.0+3202



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Swift-BAT Data (15-150 KeV)





Swift-BAT image by Stephen Kerby

Significance map for 3HWC 1912+103 and 3HWC 1914+118 (left) and the Swift-BAT observation of these sources (right). The red circles represent HAWC's field of view, and the green spots are unassociated sources from 4FGL.

Swift-BAT data for the J1928/J1930 region





Significance map for 3HWC J1928+178 and 3HWC J1930+188 (left) with their unassociated source detected by Fermi. SWIFT-BAT map (right) for these same sources with HAWC's field of view included.

3HWC J1837-066





Swift-BAT image by Stephen Kerby

Significance map for 3HWC J1837-066 (left) with a new excess found within the angular uncertainty region in SWIFT-BAT's observation (right). From our research, this was the most interesting candidate correlation that should be further looked at.

Conclusion & Outlook

- Found at least one significant association for a previous unassociated gamma-ray source
- Coincident excesses over 10 orders of magnitude in energy! From 10⁴ eV up to 10¹⁴ eV
- Study the energy spectrum of the 3HWC source from X-ray energies up to VHE gamma rays
- Write a ToO proposal for SWIFT-XRT (0.2-10 KeV)

