Optical Spectroscopy of Very Low Mass Stars

An Undergraduate Research Report



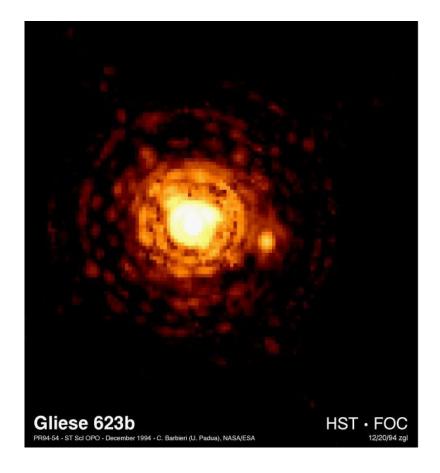


Agenda

- Introduce the Science Goals of this project
- Overview of Data and Analysis
- Preliminary Results

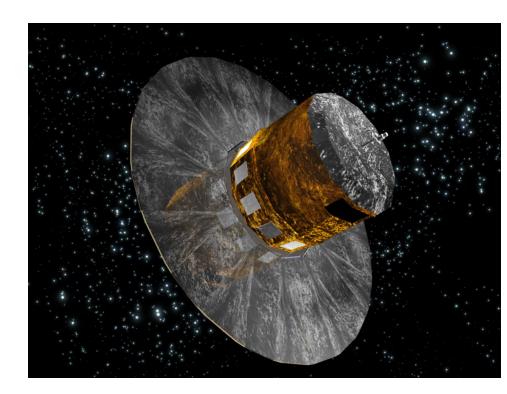
What are we looking at?

- Very Low Mass Stars and Substellar Objects
 - M type Main Sequence Stars
 - Brown Dwarfs
 - Typically <0.15 Solar Masses</p>
- Most numerous, longest living objects in the Galaxy
- They encode galactic history in their kinematics and composition



How do we Find Them?

- All-sky surveys, such as 2MASS and SDSS discovered many dim, red objects
- We are interested in stars within 30 pc of the Sun, so we can use kinematics
- Closeby stars have either small parallax or high proper motion
- Look for dim, red things that move fast!
- Our targets are selected from *Gaia* and LaTE-MoVeRS



https://www.esa.int/var/esa/storage/images/esa_multimedia/images/2006/05/artist_s_impression_of_the_gaia_satellite/9272925-5-eng-GB/Artist_s_impression_of_the_Gaia_satellite_pillars.jpg

How did We Look at Them?

The Instrument...

The Shane 3-m telescope at Lick Observatory, Mt. Hamilton, CA

The Kast Double Spectrograph





What was my Role?

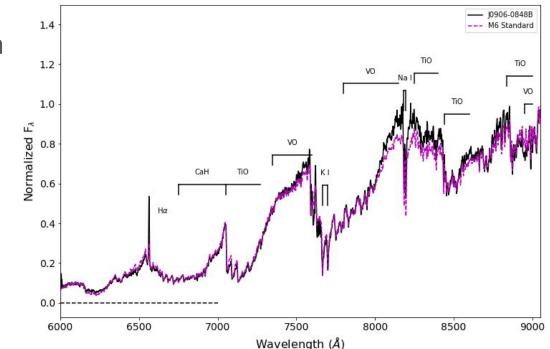
Data Reduction

Data Analysis

Data Analysis

Three major measurements:

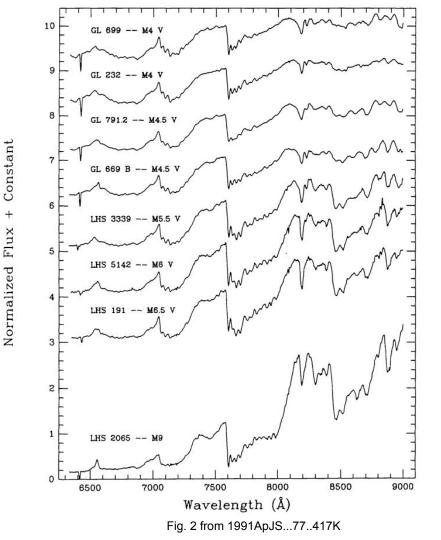
- Standard Comparison
- Spectral Indices
- Equivalent Widths



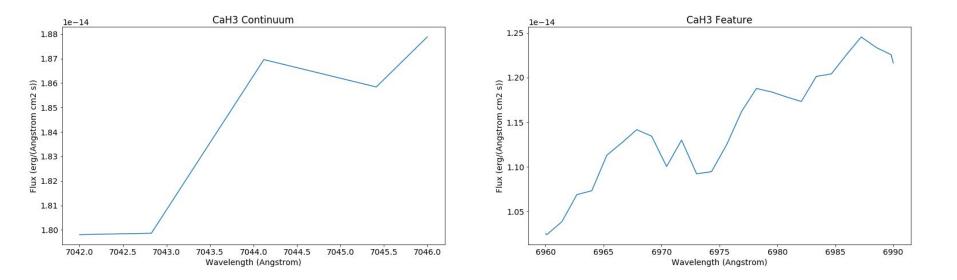
Standard Comparison

- Spectral types are defined by spectral standards
- Previous astronomers chose what stars represent each type
- One way to obtain the spectral type of some star is direct comparison to standards via chi-squared testing

$$\chi^2 = \sum_i \frac{\left(f_i - AF_i\right)^2}{\sigma_i^2}$$



Spectral Indices



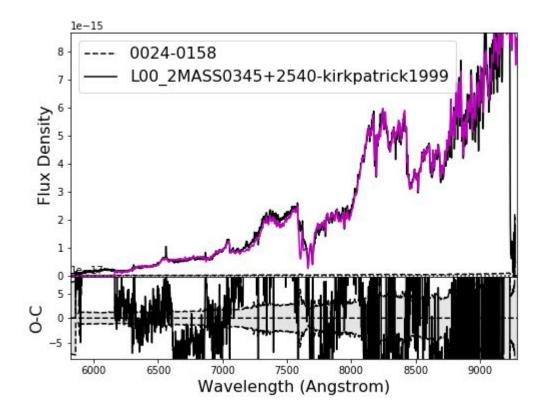
Equivalent Widths

Equivalent Width of Klb 11 1.0 Normalized Flux Density .0 .0 0.7 0.6 7680 7690 7710 7720 7730 7670 7700 Wavelength (Angstrom)

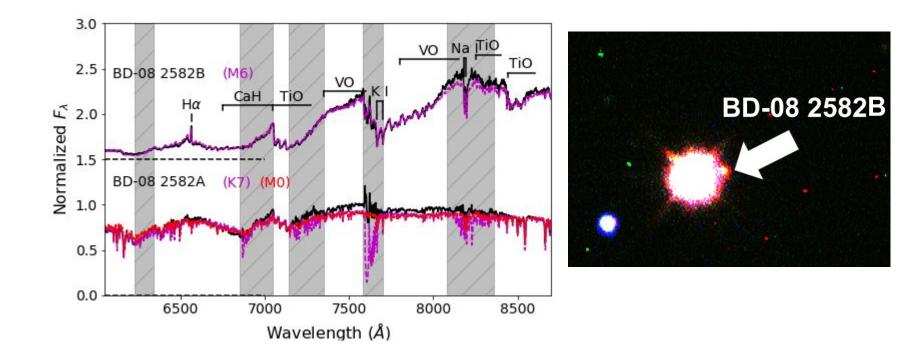
2.481673000261539 Angstrom

Preliminary Results?

Looking for "Oddballs"



Potentially Interesting Systems



Thank You!

Any Questions?